

The latest improvements in VK3UM's EMECalc software

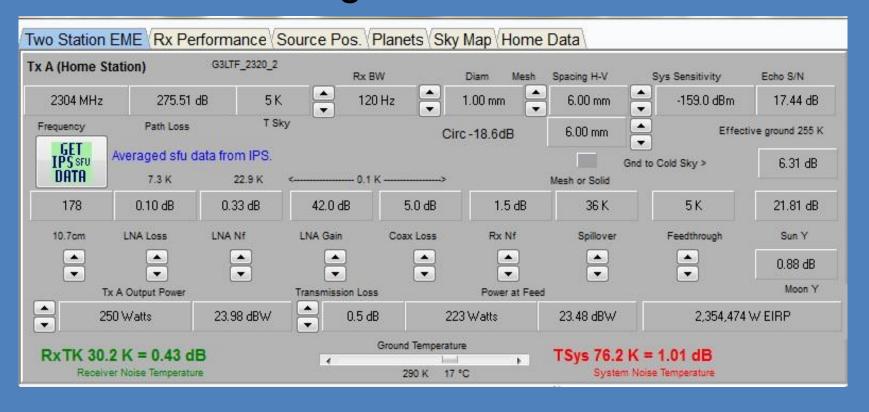
- Better Spillover estimation giving more accurate G/T value.
- More accurate Moon temperature estimate.

Components of Tsys

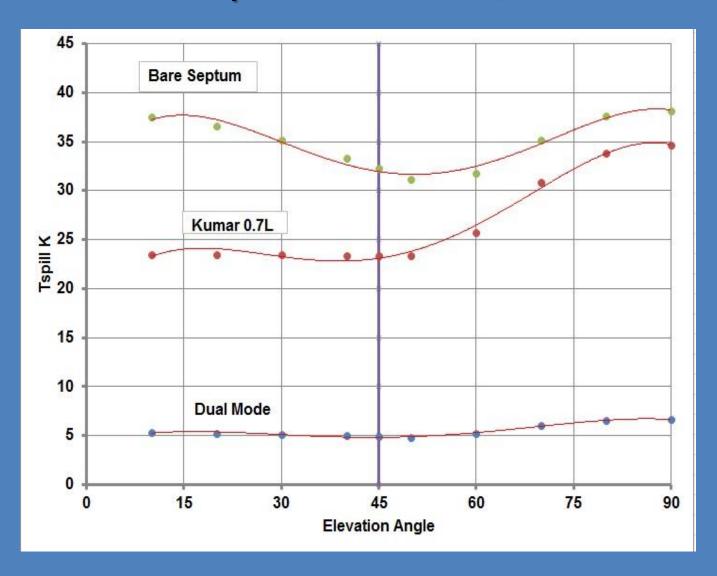
Losses ahead of the LNA LNA Noise Figure and second stage contribution

Spillover Ts

Mesh Feedthrough



Spillover contribution varies with Elevation Examples at 0.4 f/D, 290K



Using W1GHZ's online Antenna Book modelling.....

 Until Version 10, the spillover temperature at 45 deg elevation was derived for each f/D as:

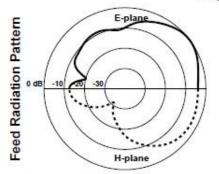
Ts=290*spillover% * 0.75

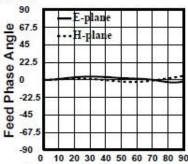
 The 0.75 was a guestimate of the amount of spillover "seeing" the ground, temperature 290K

The 0.75 is called factor Y

- Paul's EME2014 paper
 "Feedhorn analysis for parabolic dish G/T" inspired an improved method
- Dmitry, RA3AQ's paper provided the same analysis

Original VE4MA Feed 0.77λ horn diameter Ring 0.50λ wide x 0.50λ deep, 0.15λ behind rim Figure 3

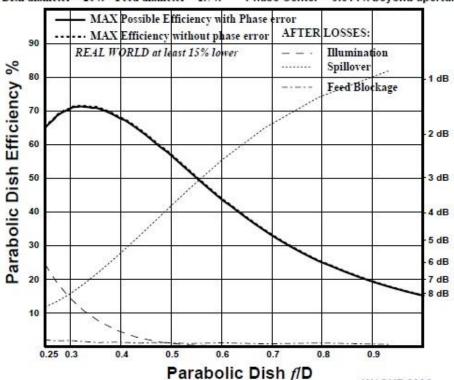




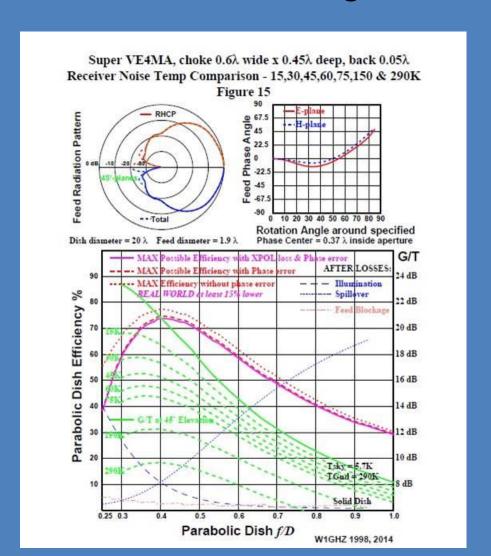
Dish diameter = 20λ Feed diameter = 1.7λ

Rotation Angle around specified Phase Center = 0.014 λ beyond aperture

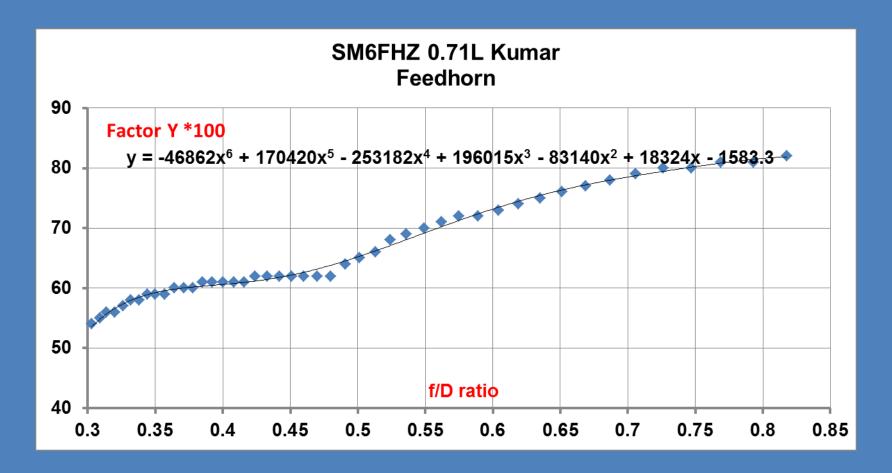
W1GHZ 2006



Curves in the W1GHZ EME2014 paper are for specific dish sizes and Trx but we can obtain a value for factor **Y** for the range of f/D ratios



A typical result for **factor Y** versus f/D at 45 deg elevation. Using this (in place of "0.75") has greatly improved the accuracy of G/T estimates



How EMECalc V10 calculates Tspill

Spillover temperature contribution Ts is calculated for each feed and f/D as:

Ts=Tg*spillover%*(6th order polynomial of Y)

Tg is ground temperature taken from the slider selected value

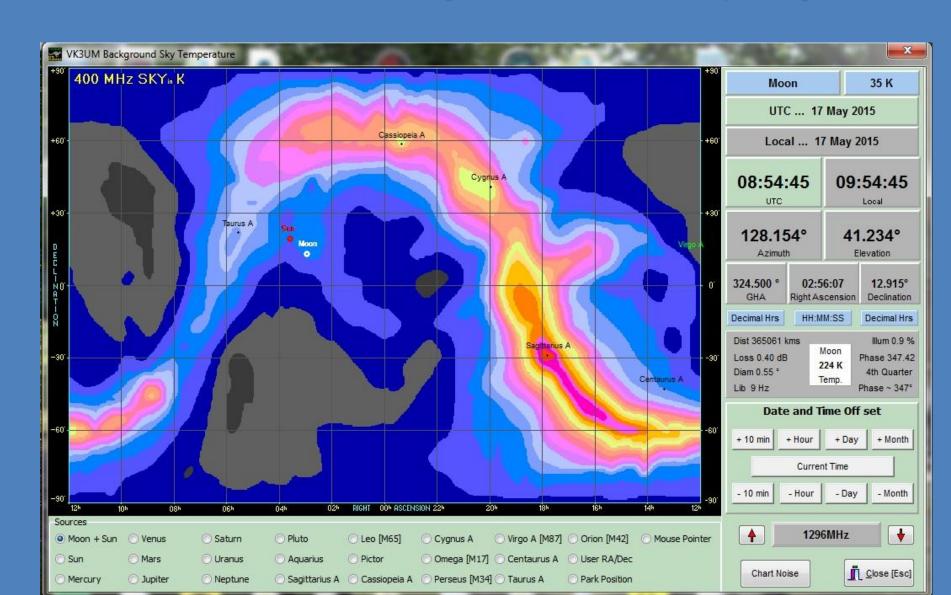
A more accurate Moon Temperature Estimate in EMECalc



Measuring G/T with Moon Noise

- To measure G/T we need two temperature reference points, one cold, one hot.
- Cold sky is relatively easy......

VK3UM - Sky noise display



Measuring G/T with Moon Noise

- To measure G/T we need two temperature reference points, one cold, one hot.
- Cold sky is relatively easy.
- Hot sources
 - Ground House wall, Tree in full leaf –must fill the near field
 - **Sun** unpredictable flux vs frequency
 - Radio stars predictable flux but weak
 - Moon higher flux but variable with distance and phase

Moon noise characteristics

 The level of moon noise received varies with frequency, moon phase and distance

At 2320MHz with 6m dish and Tsys = 73K moon noise variation 0.98 to 0.76dB

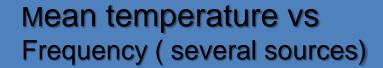
At 3400MHz with 6m dish and Tsys = 95K moon noise variation 1.45 to 1.12dB

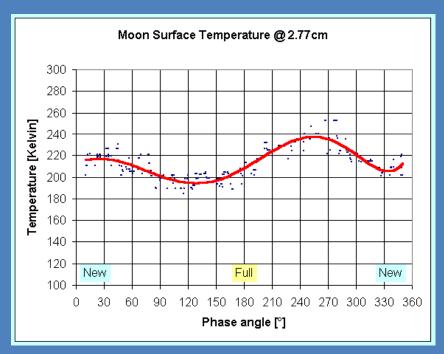
 So we need to state the conditions of our measurements.

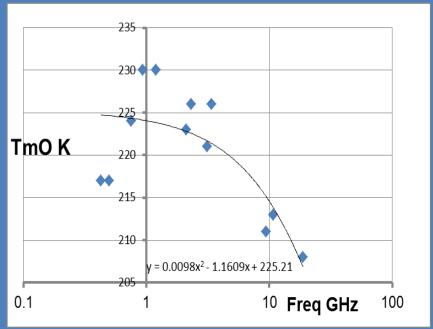
Moon Temperature vs phase and frequency

C.Monstein, Zurich 2001

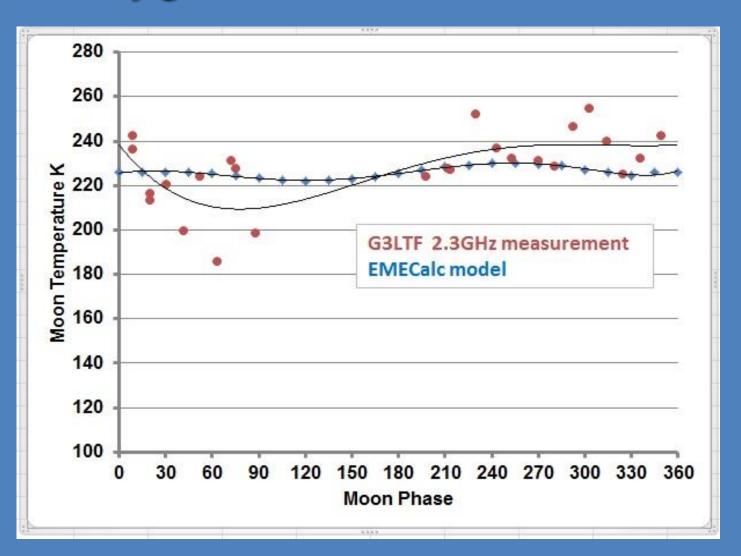
Moon temperature vs Phase



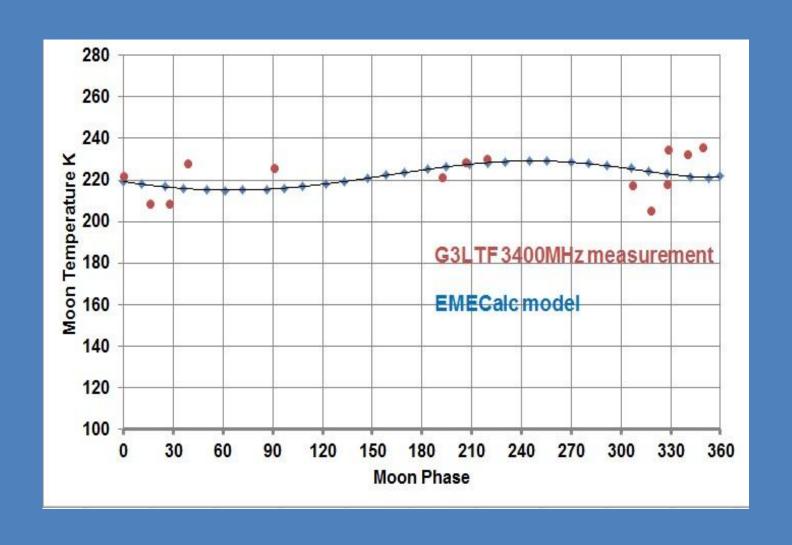




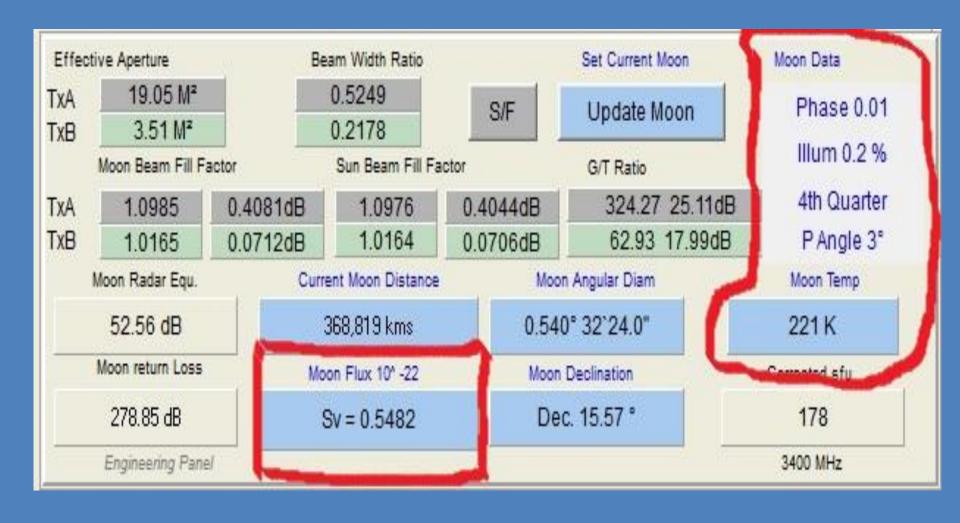
2.3GHz measurement referenced to Cygnus: Mean value 226K

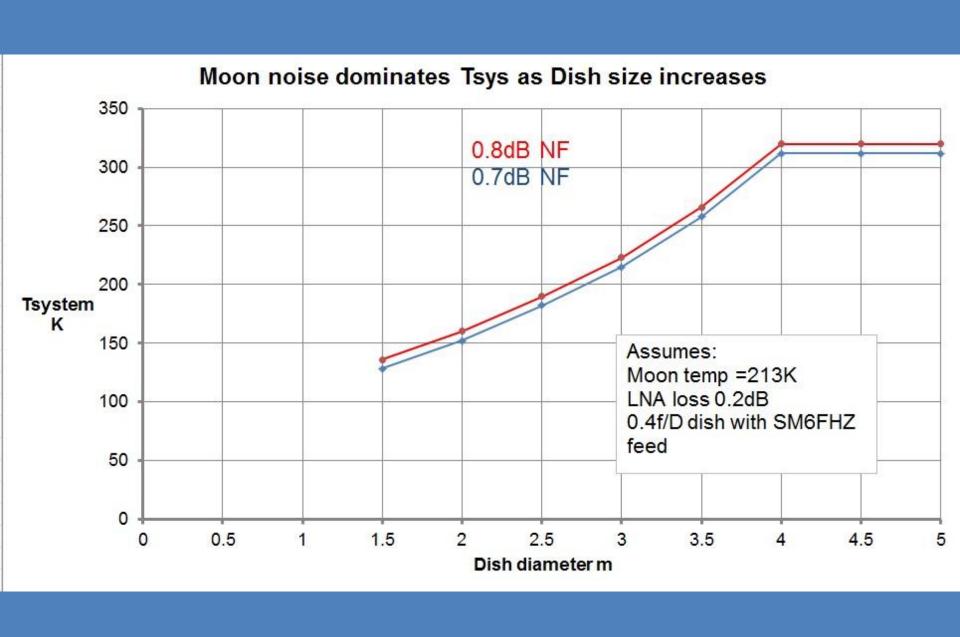


3.4GHz measurement referenced to Taurus: Mean value 223K



Moon Data in EMECalc is now better, but still work in progress.





Acknowlegements

Doug, VK3UM Paul, W1GHZ Ingolf, SM6FHZ